

Center for  
**Quantum Networks**

*NSF Engineering Research Center*



# Multi-Aperture Telescopes at the Quantum Limit of Superresolution Imaging

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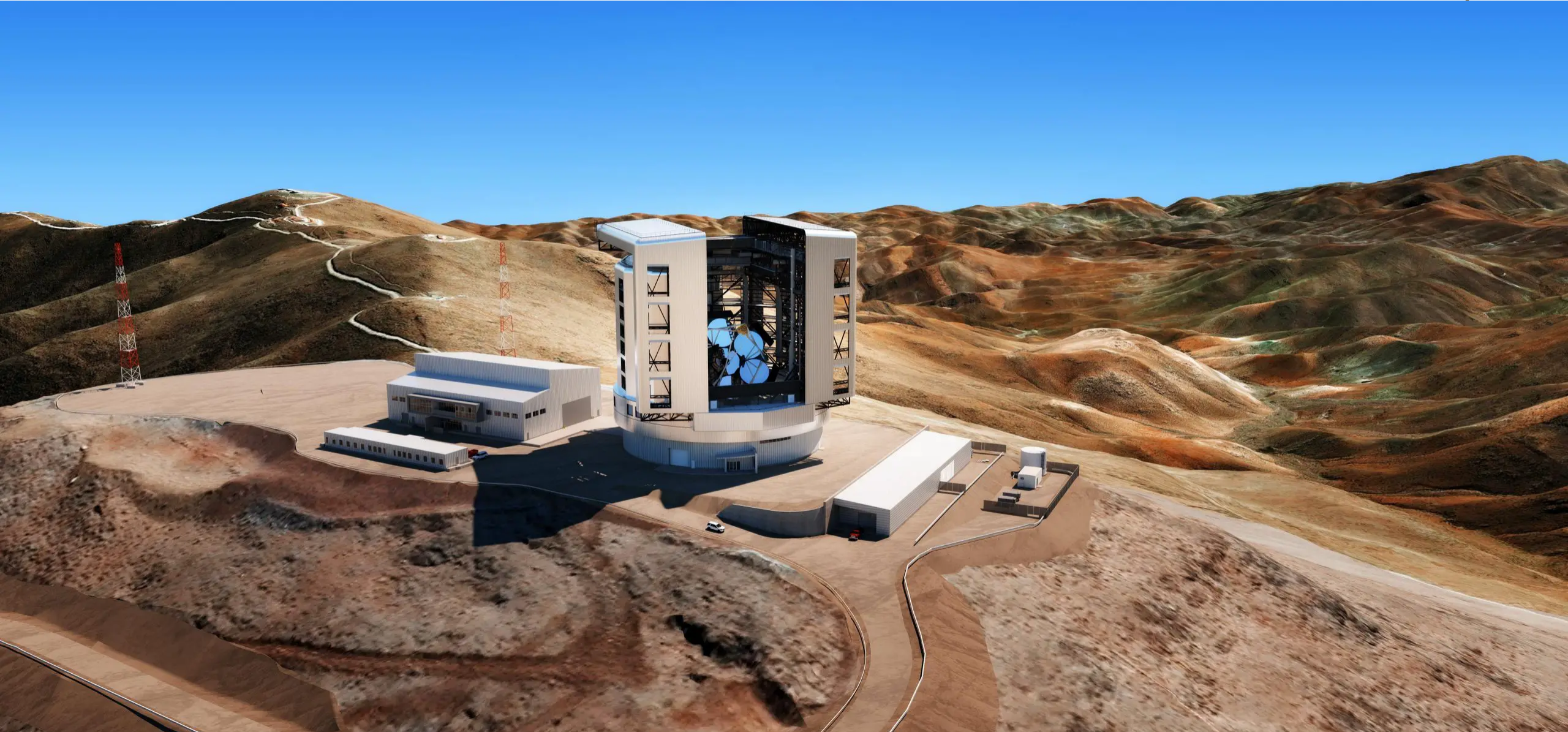
Funded by the National Science Foundation and the Department of Energy under NSF cooperative agreement #1941583



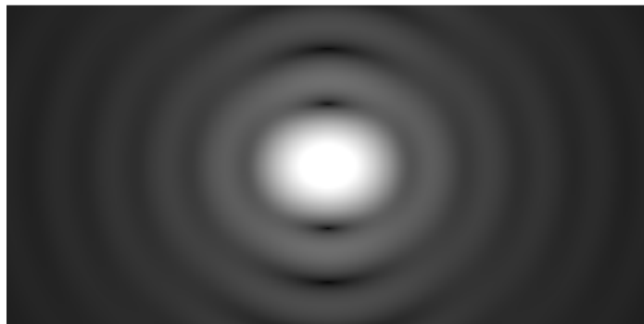
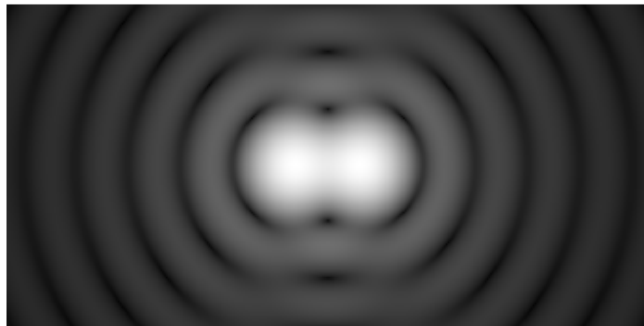
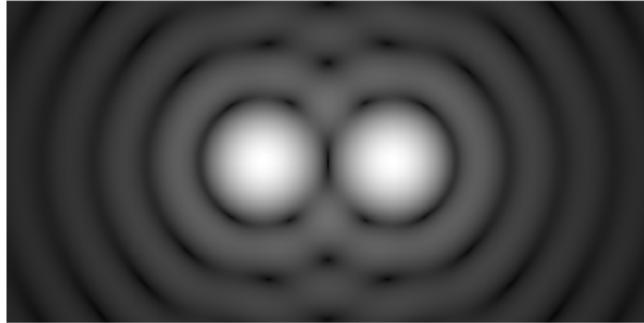
# Size Matters!



# Giant Magellan Telescope



# Rayleigh Limit



- Two point sources clearly distinguishable for a telescope of a given aperture
- Two sources at the Rayleigh limit
- Two sources beyond the Rayleigh limit are difficult or impossible to distinguish

# Rayleigh Limit

Angular Resolution

$$\theta \approx 1.22 \frac{\lambda}{D}$$

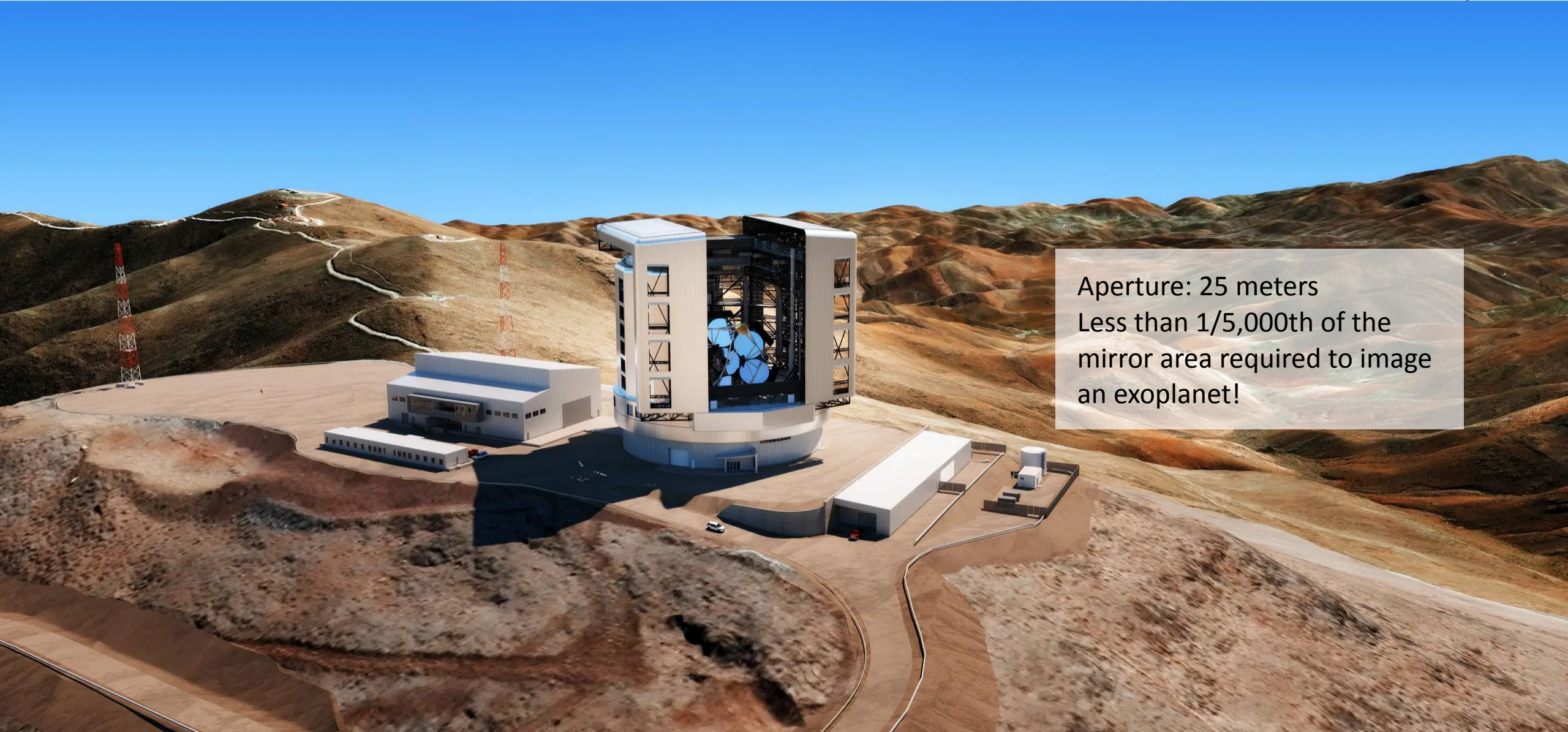
Wavelength

Diameter of mirror

**The larger the telescope mirror, the more precise the resolution... *but***

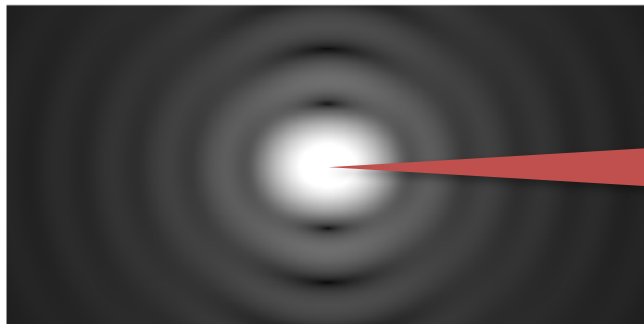
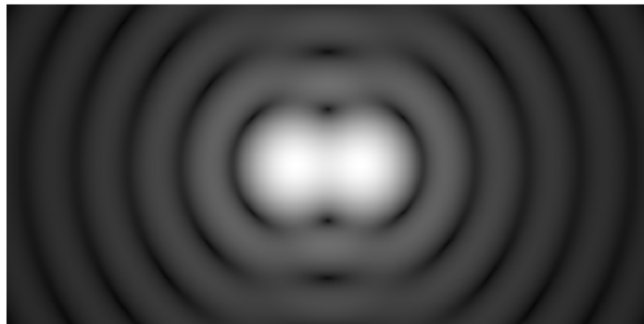
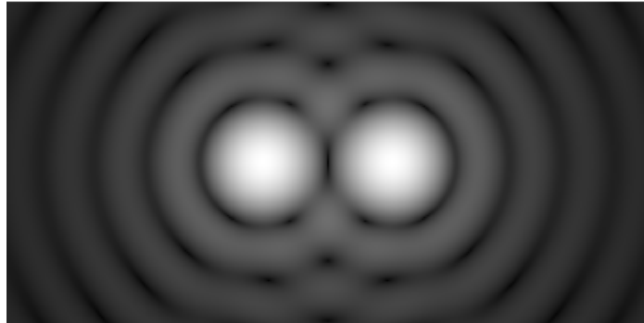
**Imaging a Earth-size planet at Alpha Centauri at 40x40 pixels would require an optical telescope with a 1.8 kilometer aperture!**

# Giant Magellan Telescope



Aperture: 25 meters  
Less than 1/5,000th of the  
mirror area required to image  
an exoplanet!

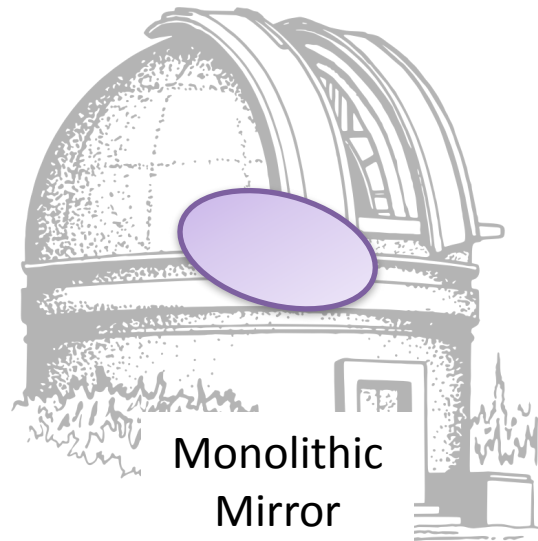
# Rayleigh Limit



- Two point sources clearly distinguishable for a telescope of a given aperture
- Two sources at the Rayleigh limit

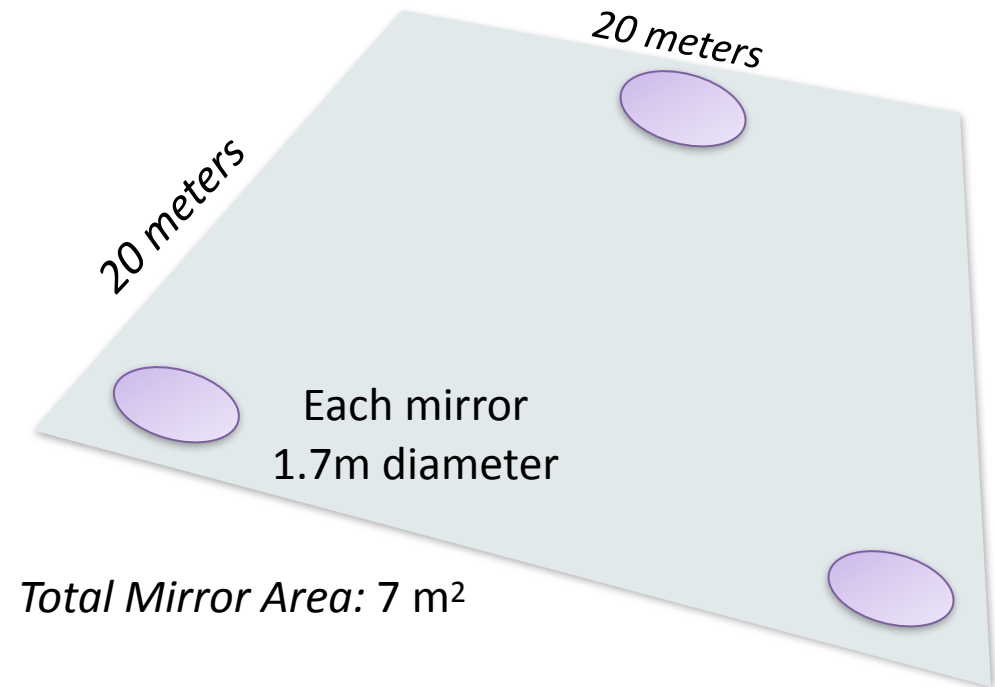
- Two sources are too close to be distinguished with a classical telescope. Quantum techniques can beat the Rayleigh limit by unlocking all the information about amplitude and phase in the collected light.

# Multi-Aperture Telescopes



Monolithic  
Mirror

3m diameter  
Area: 7 m<sup>2</sup>

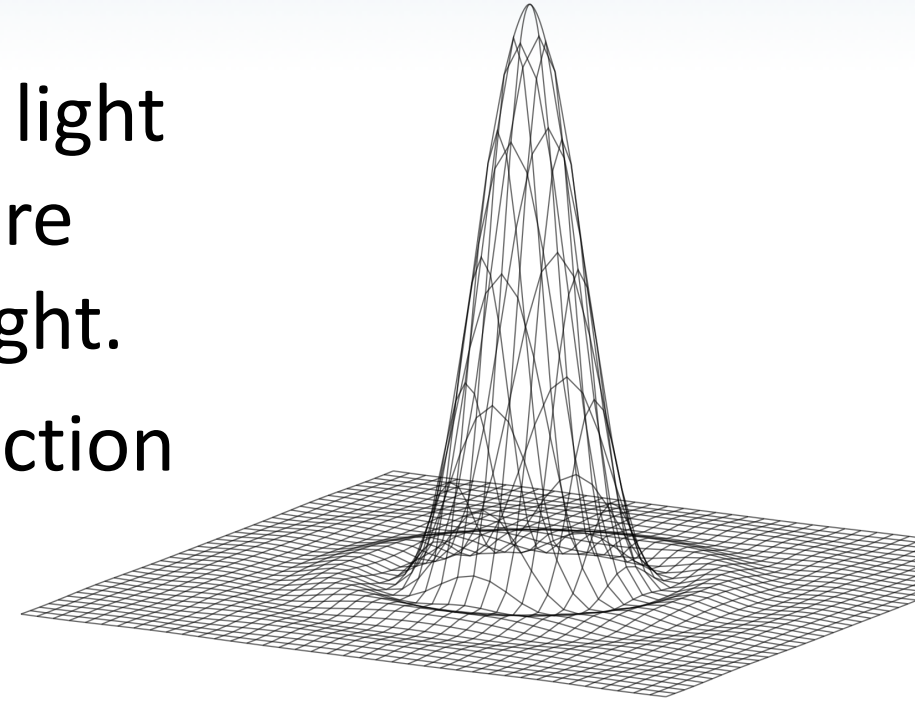
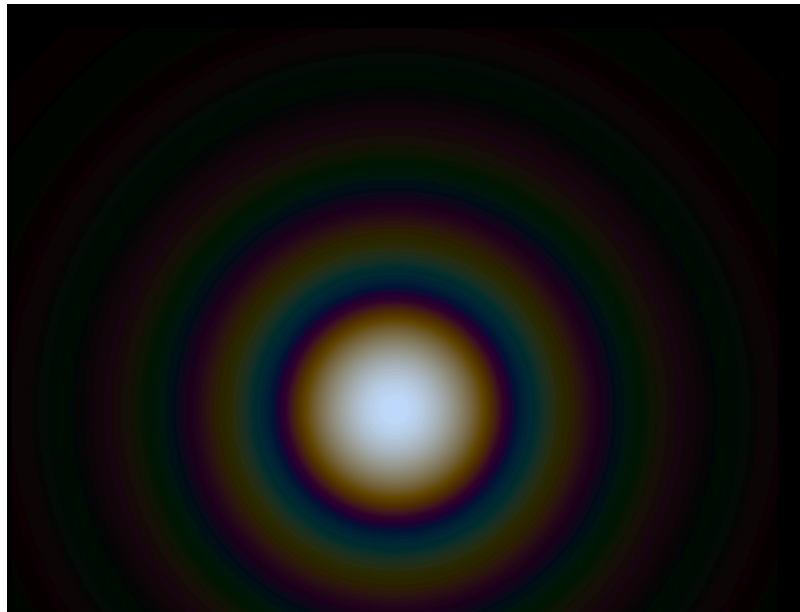


**Three 1.7m mirrors are much cheaper than one 3-meter mirror!**



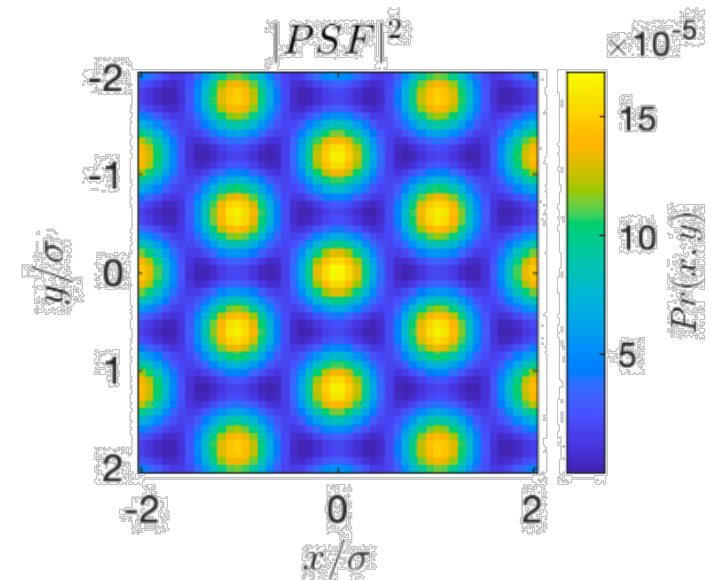
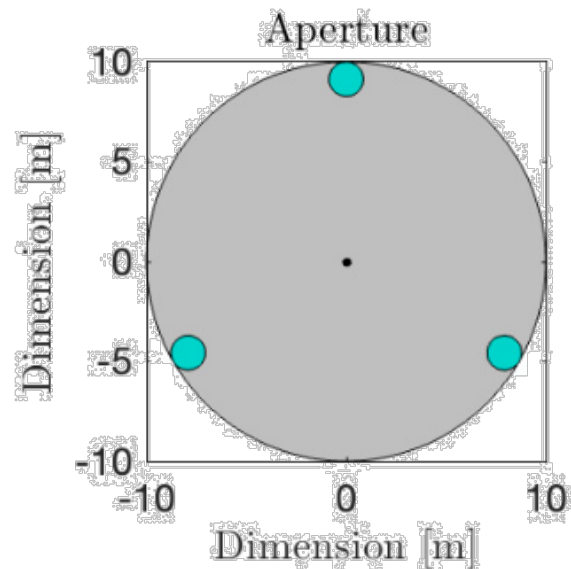
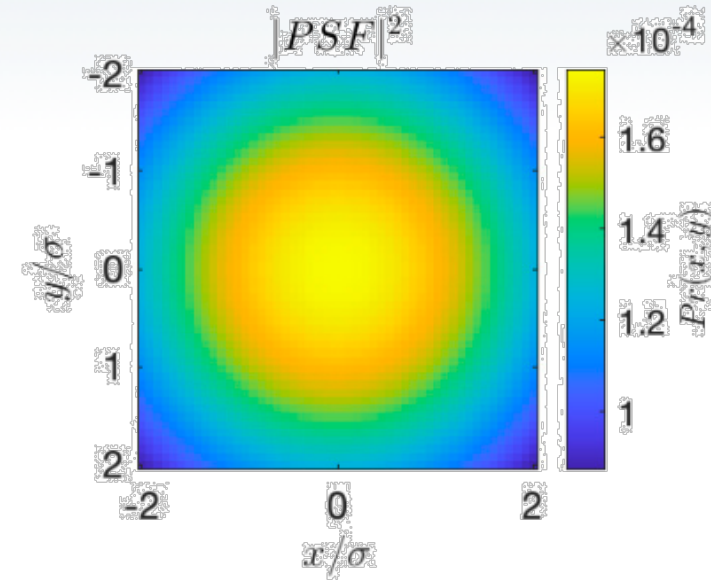
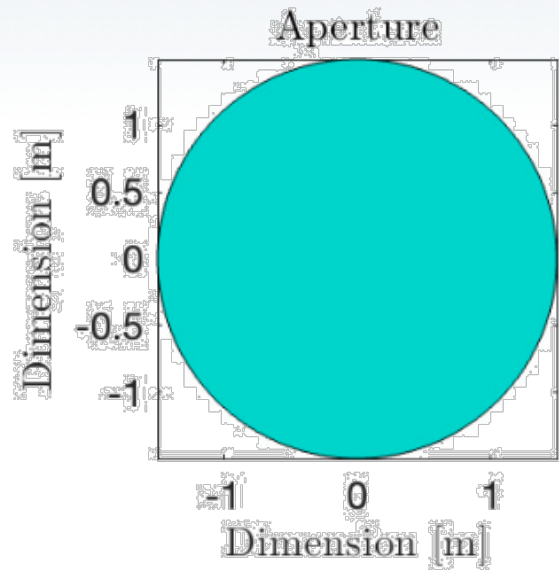
# Airy Disk

- The Airy Disk is the best-focused spot of light that a perfect lens with a circular aperture can make, limited by the diffraction of light.
- This is an example of a Point Spread Function (PSF).

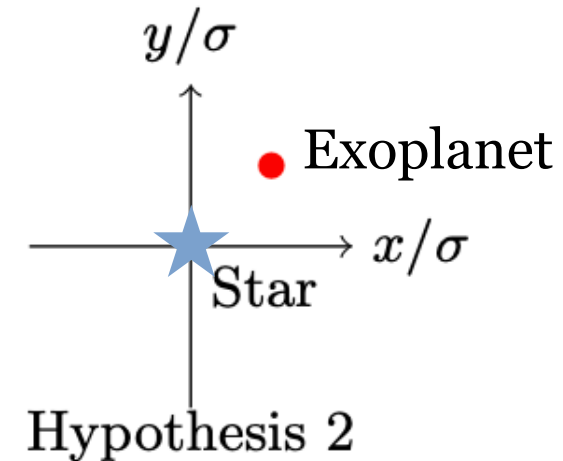
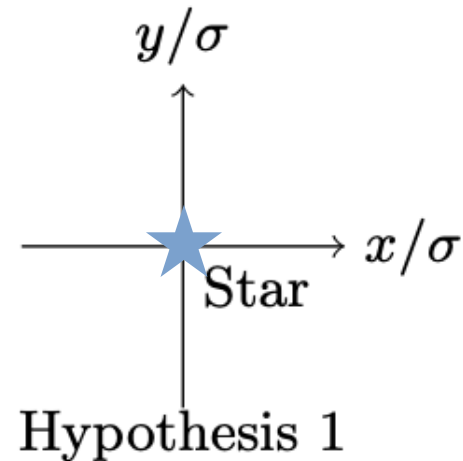


Different wavelengths will diffract differently

# Point Spread Functions

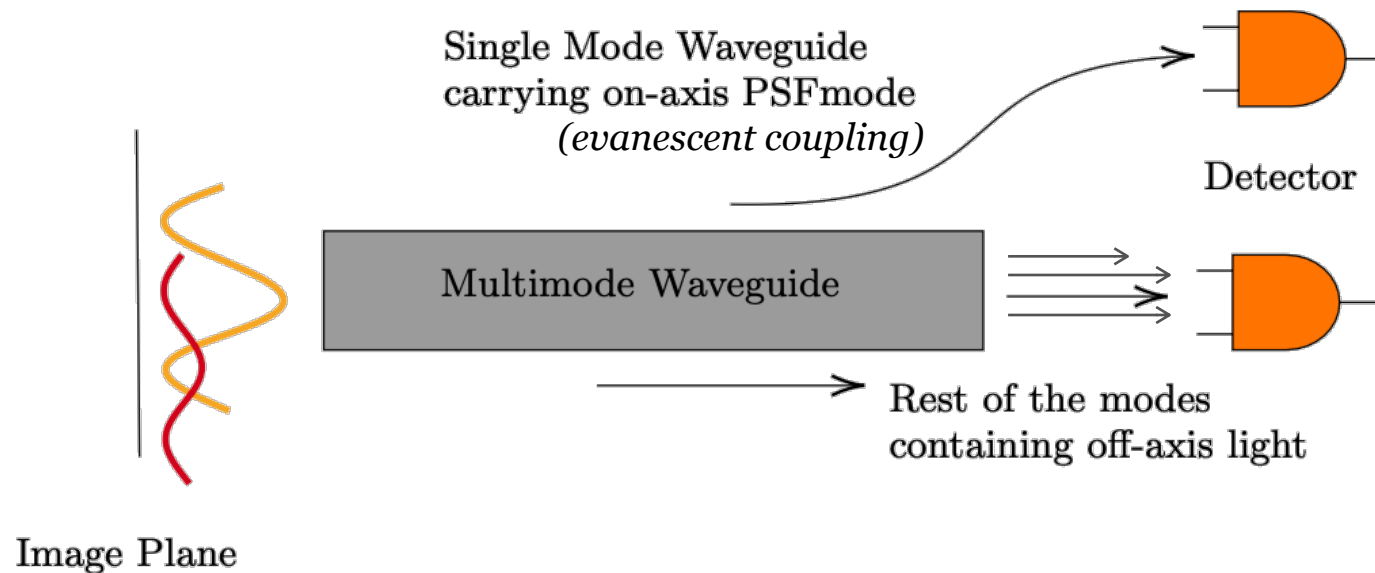
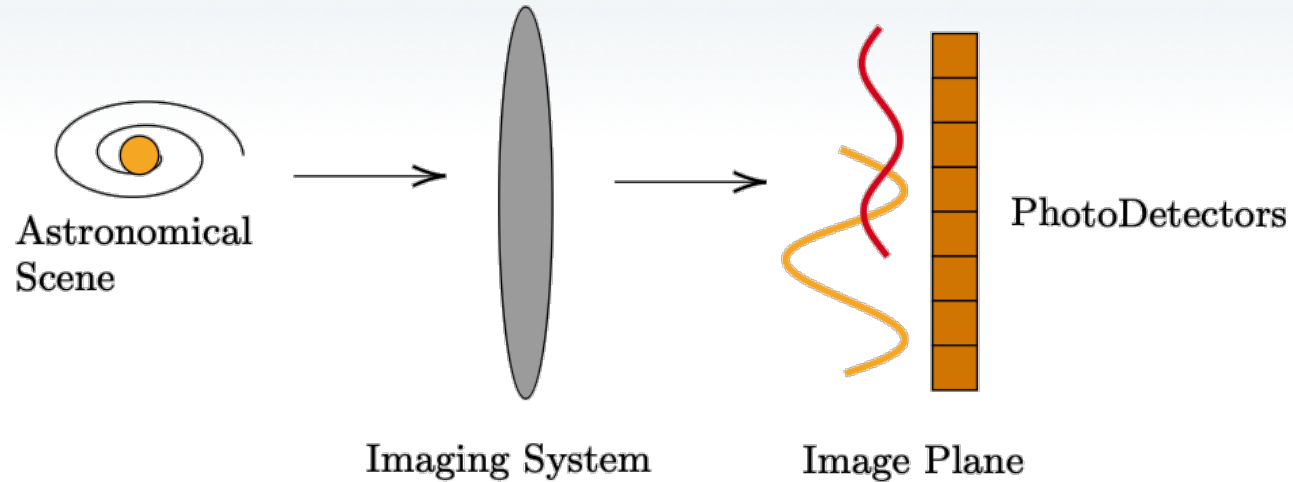


# Resolving an Exoplanet

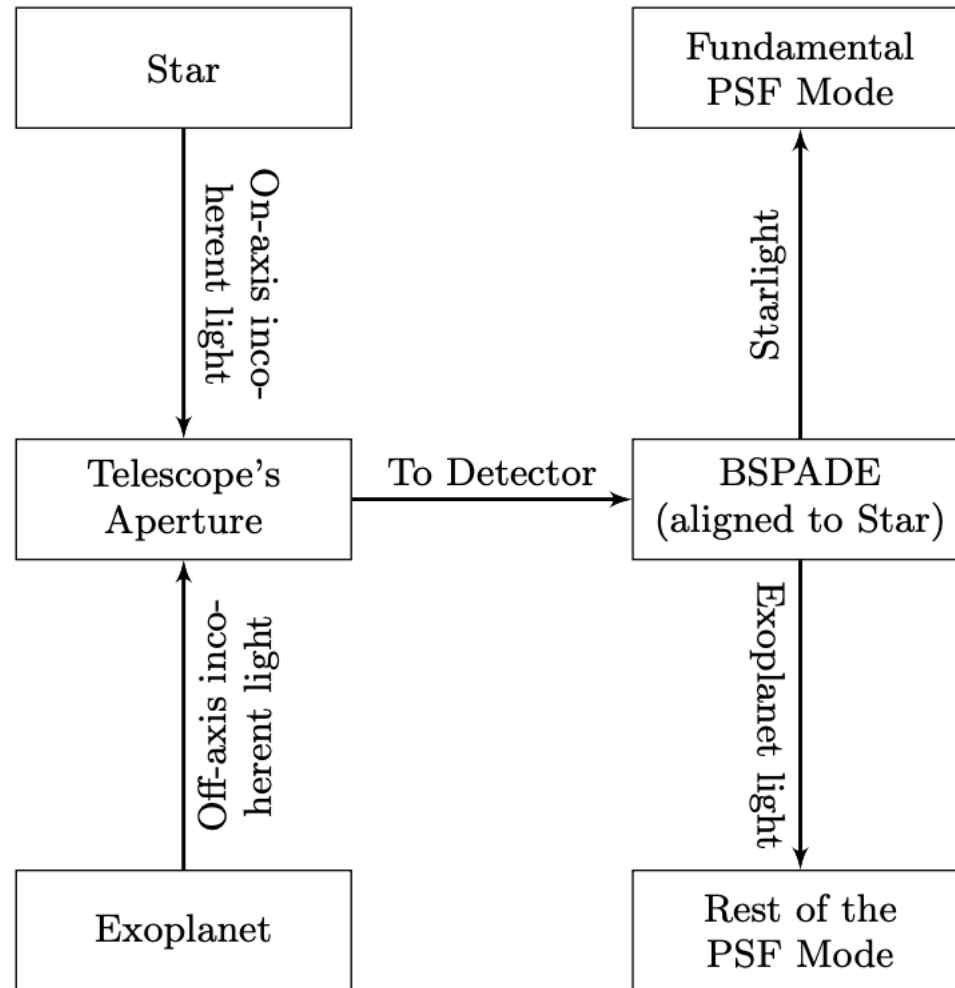


**First test: Is a second source of light (exoplanet) present or not?**

# Direct Imaging vs. BSPADE



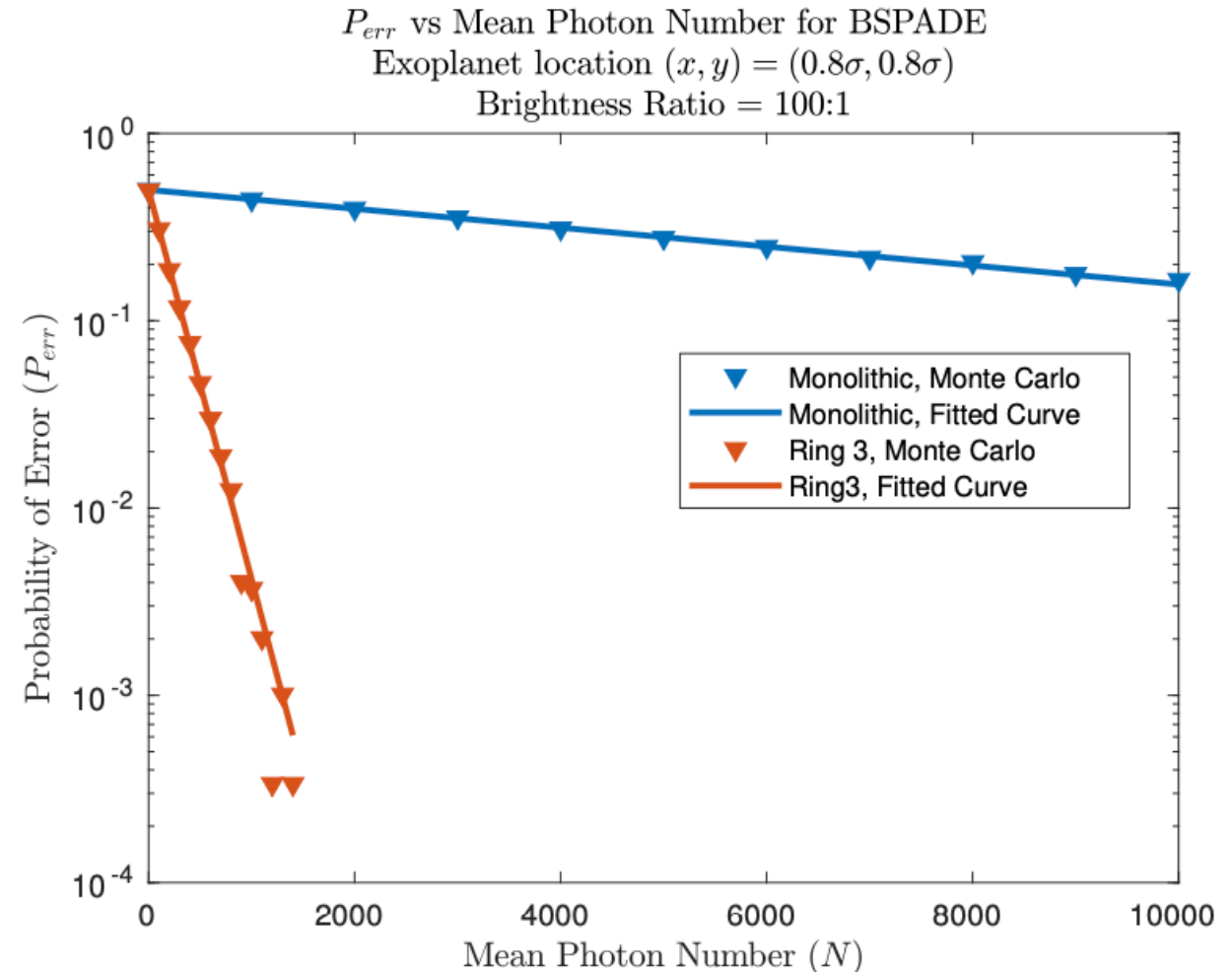
# BSPADE Detection



- BSPADE: Binary Spatial Mode Demultiplexing
- Telescope(s) are aligned with Point Spread Function (PSF) central star on-axis
- Off-axis photons are sorted into separate detector — these are (mostly) from the exoplanet(s)
- Acts as “nuller” for starlight

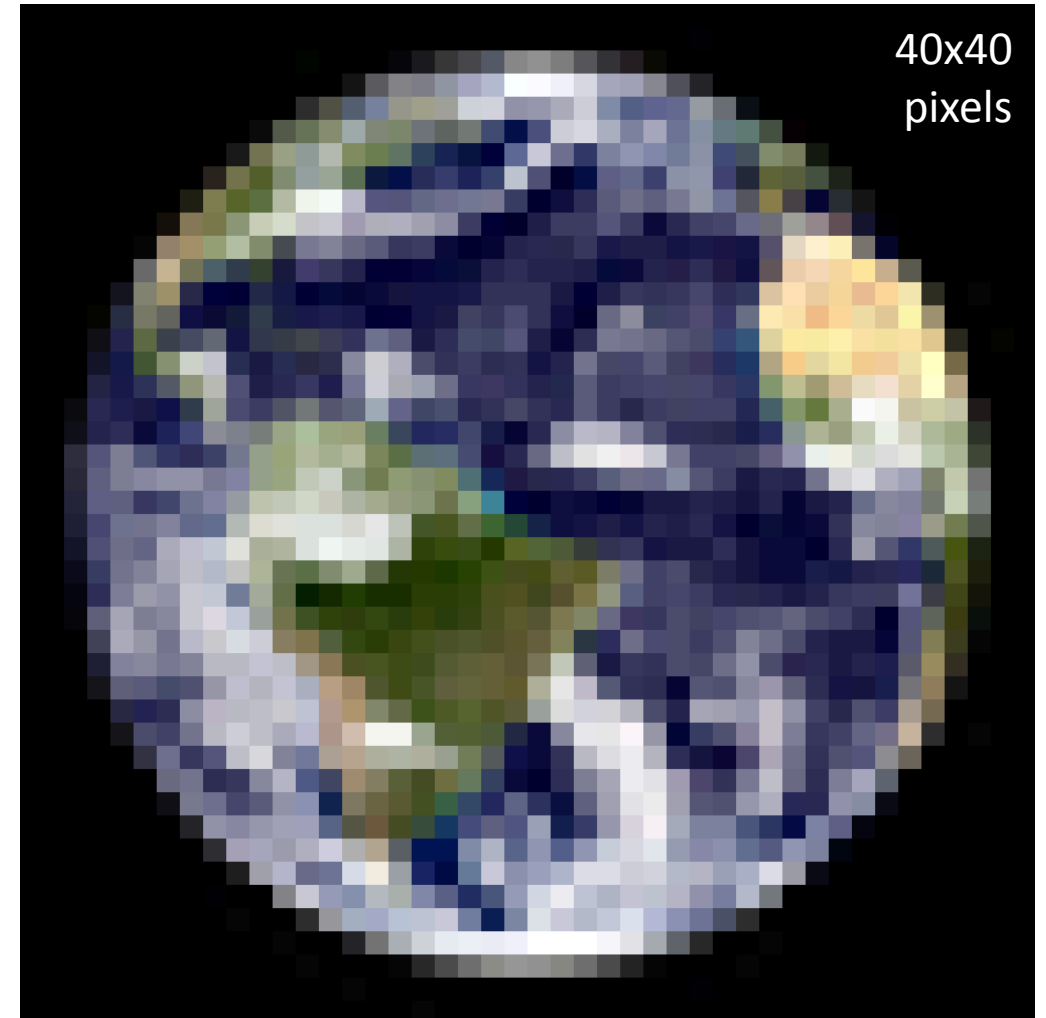
# Probability of Error Simulation

- Multi-aperture arrangement (Ring3) substantially outperforms monolithic aperture
- Substantially increased probability of distinguishing exoplanet from star, even with limited photon flux



# Conclusions / Future Work

- 100x increase in sensitivity for multi-aperture compared to monolithic mirrors for identical image-gathering area
- Ability to use BSPADE as starlight nuller in detecting and discriminating light from exoplanet(s)
- *Future development:* using off-axis photons to go beyond detection of exoplanets to direct imaging



# Thank You!



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